# SOTERIA

#### Giovanni Lapenta for the Soteria Consortium

Centrum voor Plasma-Astrofysica Katholieke Universiteit Leuven BELGIUM

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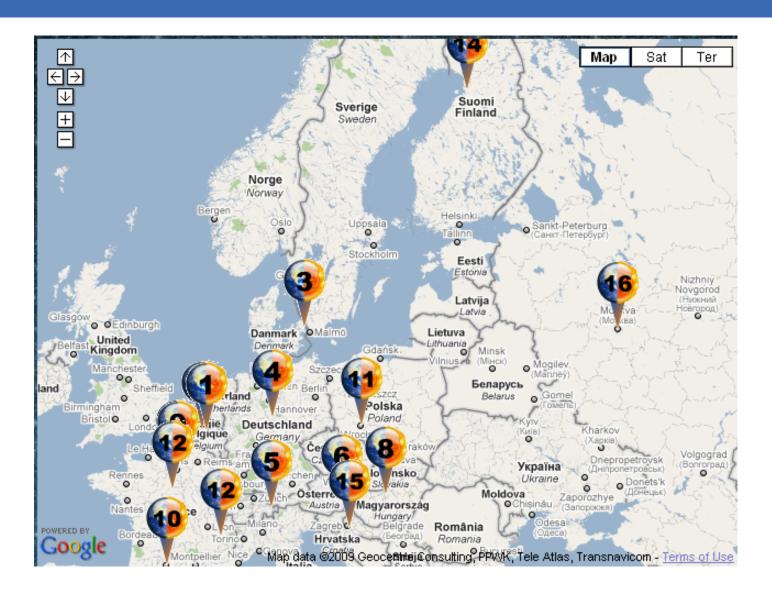
### **Outiline**



- Overview of the project
- Overview of the progress
- Future events

## Map of Soteria





# **SOTERIA EC network funded by the EC/FP7**







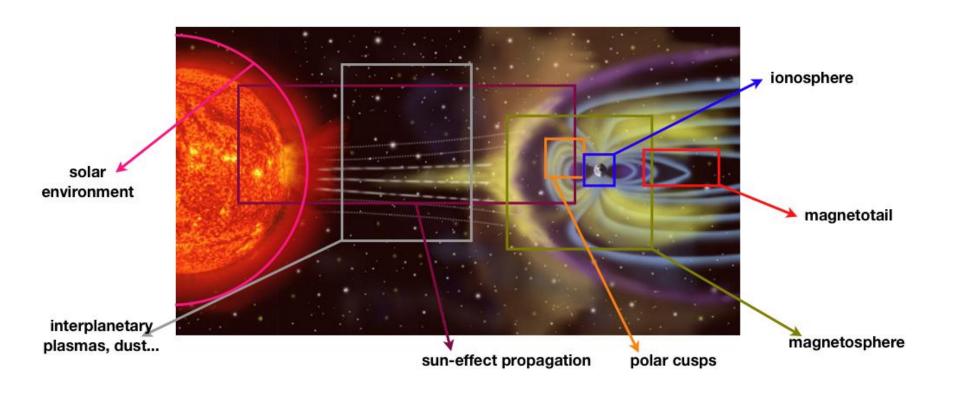
- Coordinator: G. Lapenta
- Data (ground & space) and simulation on:
- 1. Photosphere
- 2. Chromosphere/Corona
- 3. Heliosphere/Terrestrial effects
- 4. Irradiance
- Focus on data dissemination
- Looking forward to collaborations outside the consortium



			LEUV
Participant Number	Participant short name	Participant organisation name	Country
(coordinator)	KU Leuven	Katholieke Universiteit Leuven	Belgium
2	UNIGRAZ	Universitaet Graz	Austria
3	PMOD-WRC	Pyhsikalisch-Meteoroligisches Observatorium Davos and World Radiation Center	Switzerland
4	КО	Konkoly Observatory	Hungary
5	CNRS LPCE & LP	Centre National de la Recherche Scientifique	France
6	ROB/SIDC	Koninklijke Sterrenwacht van Belgie	Belgium
7	OBSPARIS	Observatoire de Paris	France
8	SRC-PAS	Space Research Centre, Polish Academy of Sciences	Poland
9	MTA-KFKI- RMKI	MTA-KFKI-RMKI Research Institute for Particle and Nuclear Physics	Hungary
10	DTU	Technical University of Denmark	Denmark
11	UOulu	University of Oulu	Finland
12	UGOE	Georg-August-Universität Göttingen Stiftung Öffentlichen Rechts	Germany
13	HVAR	Hvar Observatory, Faculty of Geodesy, University of Zagreb	Croatia
14	NOVELTIS	Noveltis Sas	France
15	FIAN	P.N. Lebedev Physical Institute	Russia
16	IEEA	Informatique Electromagnetisme Electronique Analyse numérique	France

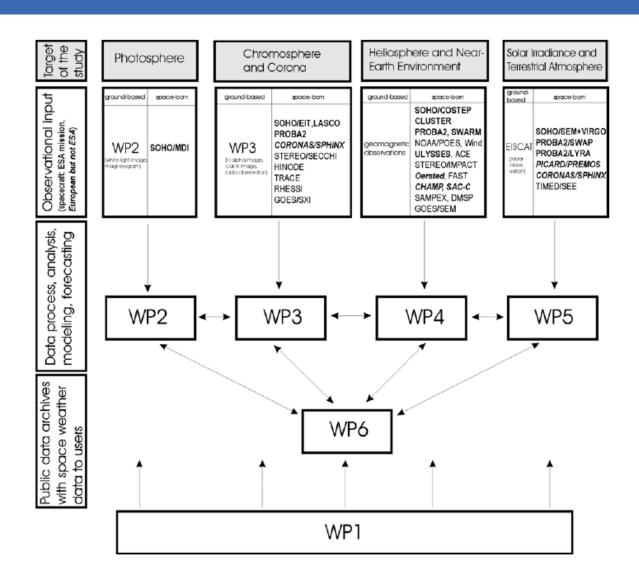
# Overview of the Space Covered by Soteria





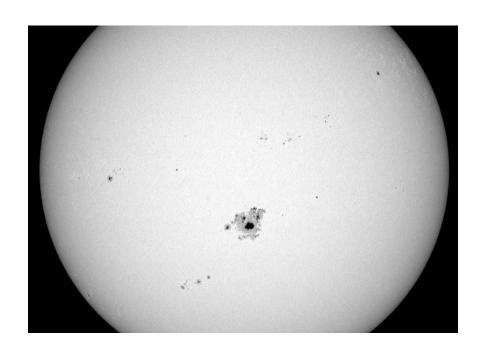
# WP1: Management Goals of the project





### **WP2: Photosphere**





Major sunspot group passing central meridian: this image shows an active Sun, only two years after the last maximum of activity, with a major sunspot group passing central meridian. This whole-disk CCD image was taken on Aug. 17, 2002 with the Uccle solar telescopes (ROB, Brussels), one of the ground-based synoptic instruments that will support Work Package 2 of the SOTERIA project.

# Relevance to space weather:

Active regions

Sunspot size, position, polarities and trends

Emerged field structures

Unstable configurations and current layers

Magnetic reconnection and initiation of CMEs

Relevance to irradiance (weather, climate)

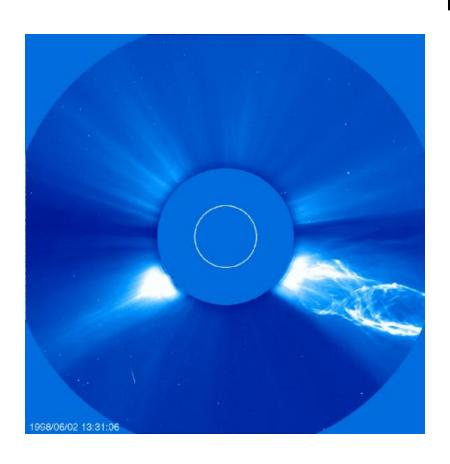
Positive: faculae

Negative: sunspots

# WP3: Chromosphere and corona





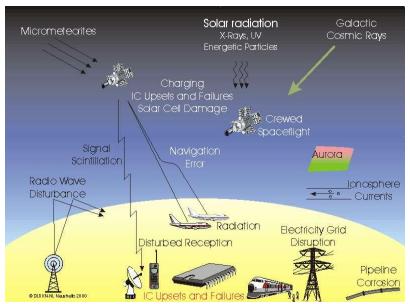


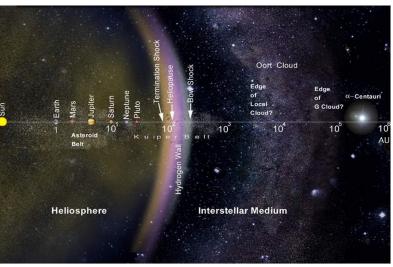
#### Focus on:

- Flares
- CMEs
- Coronal heating
- Wind creation in coronal holes (fast)
- Wind creation in helmet streamers (slow)
- Constraining models with observations

# WP4: Heliosphere and Terrestrial Effects



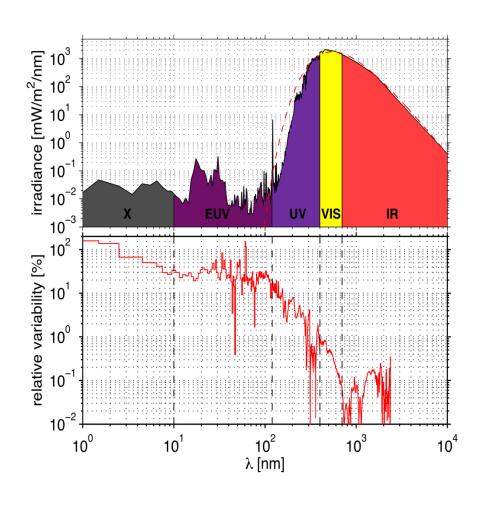




- Solar effects on the Earth, other planets, human activities
- Energetic particles streaming from the Sun
- Magnetic storms
- Variability and interactions with magnotospheres
- Processes initiated in the magnetospheres (e.g. substorms)

## WP5 Irradiance: Background





How does the **solar irradiance** vary in the EUV-UV-visible range?

What causes the irradiance to change on short (days) to long (years) time scales?

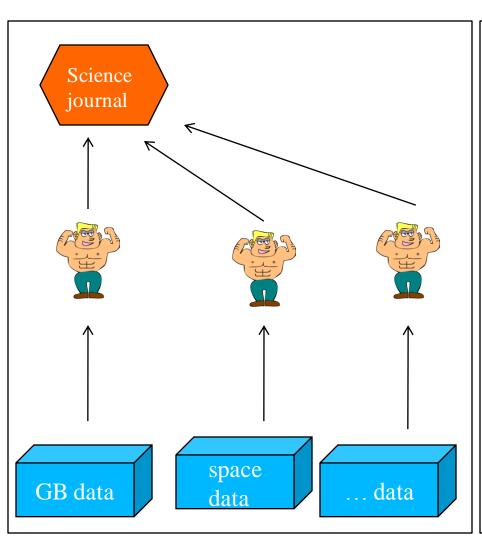
How does this affect the upper terrestrial atmosphere?

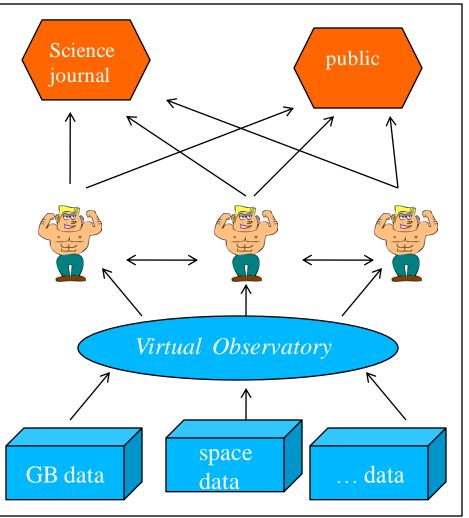
### WP6: Distribution of data and dissemination



#### **BEFORE SOTERIA**

#### **SOTERIA**





## Overview of the progress of each WP



**WP1: Management (Lapenta)** 

**WP2: Photosphere: Dr. – Ludmany** 

WP3: Chromosphere and Corona: Dr. Bothmer

WP4: Heliosphere and Terrestrial Effects: Dr.

**Vennerstrom** 

WP5: Irradiance: Dr. Dudok de Wit

WP6: Distribution of and dissemination: Dr. Berghmans

### Soteria Web Page

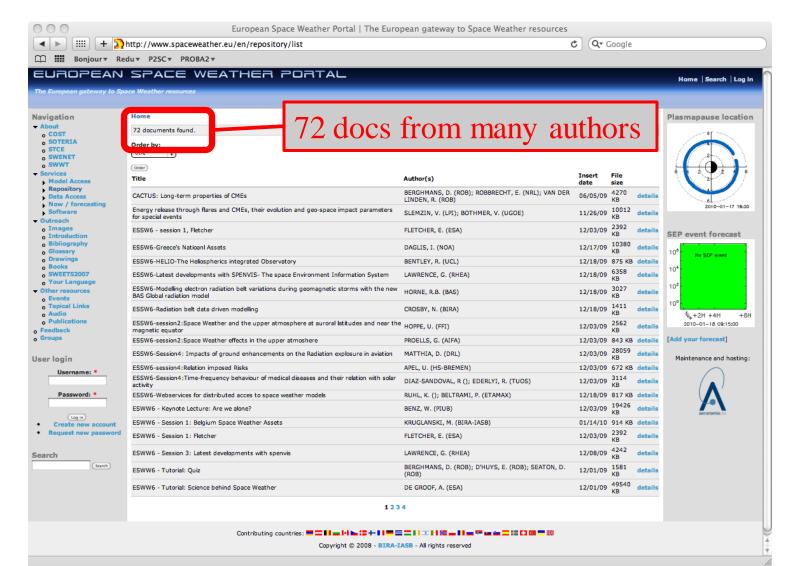




## **Document Repository**

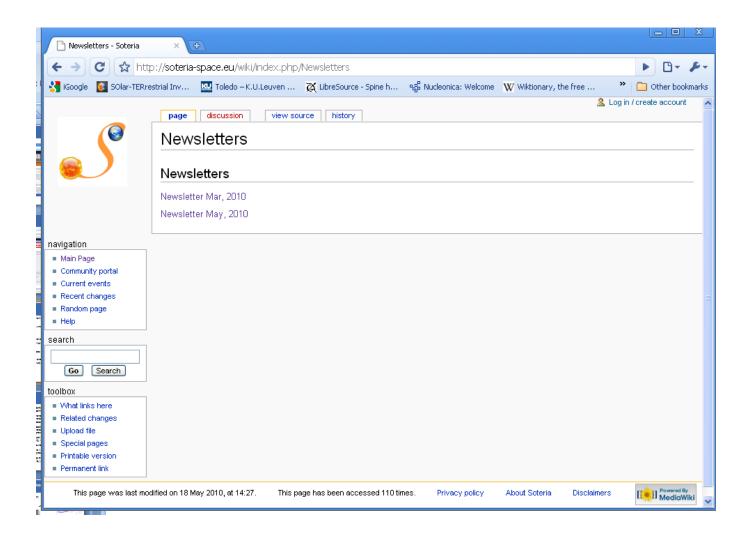


## Document repository at <a href="http://www.spaceweather.eu">http://www.spaceweather.eu</a>



#### **Newsletters**





#### **WP2 - Successes**



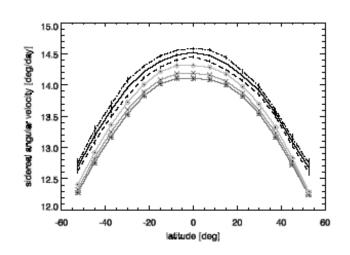


Fig. 1. Rotation velocity derived from ring-diagram analysis at 3Mm (asterisk), 7Mm (cross), 15Mm (diamond) (lines from dark to light grey) and SBCS tracing (black) averaged over the periods: August 2001 – March 2004 (dashed), April 2004 – December 2006 (dashed dotted).

Zaatri et al. (Brajsa) 2009 *A&A*, **504**, 589.

- Vibrant collaborations established among the observatories
- •Historical data, example of summer student at ROB
- •Online repositories to be linked with the VO (list on web site), examples:
  - •SOHO/MDI Continuum faculae: produced by schedule
  - SOHO/MDI Sunspot Data (SDD): produced by schedule
  - daily magnetic observations
  - •Debrecen Photoheliographic Data sunspot catalogue

#### Publications

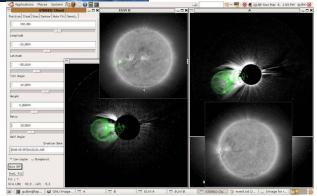
- -differential rotation (published, see figure)
- -torsional oscillation of hemispheric cycles (submitted)
- -active longitudes cooperative study, in progress
- -comparative analysis of DPD, SDD and ISN cooperative study, in progress

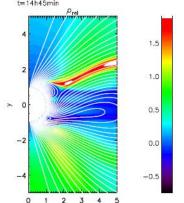
#### **WP3 - Successes**

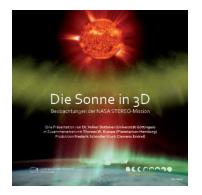












- •STEREO operations and archiving smoothly proceeding.
- •Coronas/Photon successfully launched and operated. TESIS and SphinX instruments databases established.
- •Proba2 successfully launched and operated. SWAP and LYRA database in progress.
- •Report on 3D structure (STEREO) and energy release of flares and CMEs.
- Ground-Based Telescope Developments
- •Data Analysis Modelling, Simulations in Progress.
- E/PO Material Established (DVD, bluray).

# WP4: Progress so far - Heliosphere



Carrington Rotation 2026:0 - 2027:5 Source Surface Radius: 2.50 Central Meridian Carrington Longitude 0

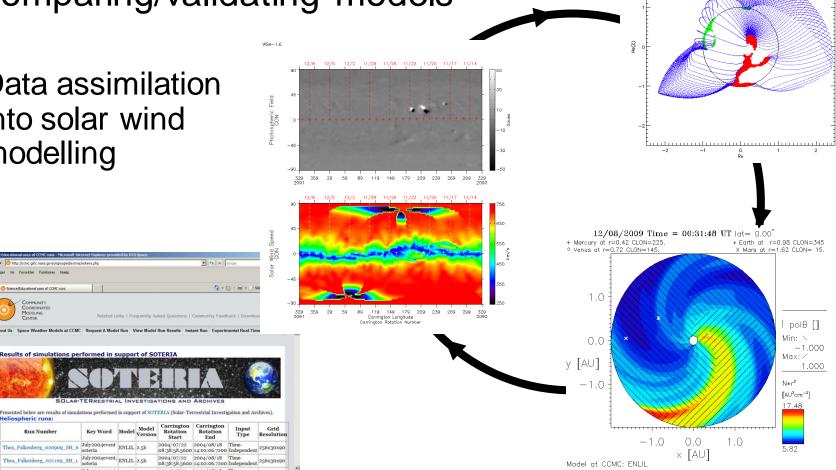
Soalar Wind, CME propagation: Comparing/validating models

End

Data assimilation into solar wind modelling

Results of simulations performed in support of SOTERIA

July2004event ENLIL 2.5b



Negative Open

Closed Field Lines

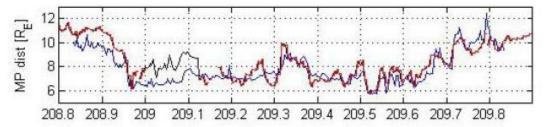
Positive Open Field Line Footpoints

## WP4: Progress so far - Magnetosphere

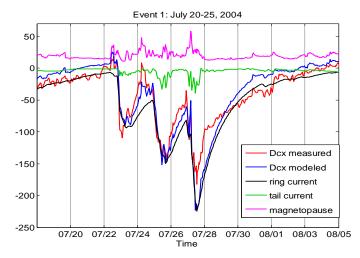


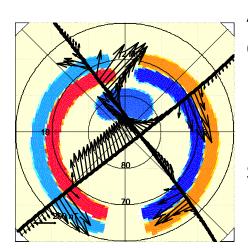
Magnetopause location: Validating and comparing models

# Coronal holes - Predicting Dst



# Geomagnetic monitoring: Dst/Dcx – separating contributions



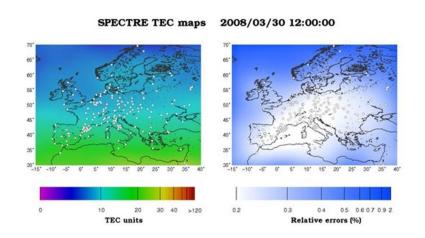


Auroral oval currents –

Inversion of magnetic satellite data

## WP4: Progress so far - Ionosphere



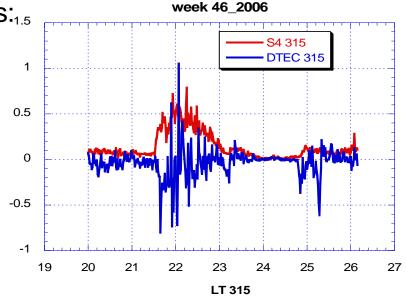


Provision of TEC Maps and timeseries from GPS data

Comparison between focus parameters: ... Scintillations, TEC and Dst

Evidence for relation between scintillations and TEC rate of change.ជ

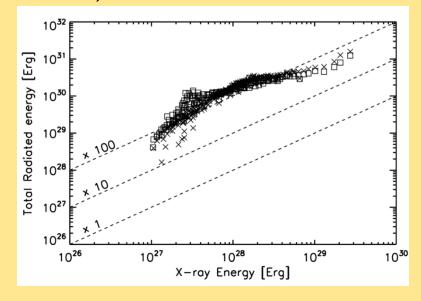
No correlation found between Dst and low latitude scintillations



## WP5 Irradiance: Progress



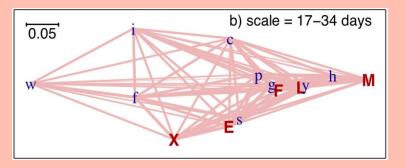
All solar flares emit most of their energy in visible light (Kretzschmar et al., 2009)



#### **New directions:**

- \* Long-term variability of the total solar irradiance: modelling and implications on climate change
- \* Response to ESA Space Situational Awareness programme: input to orbitography models
- \* Workshop on future trends in solar spectral irradiance measurements

How well do various solar proxies describe spectral bands? (Dudok de Wit et al., 2009)

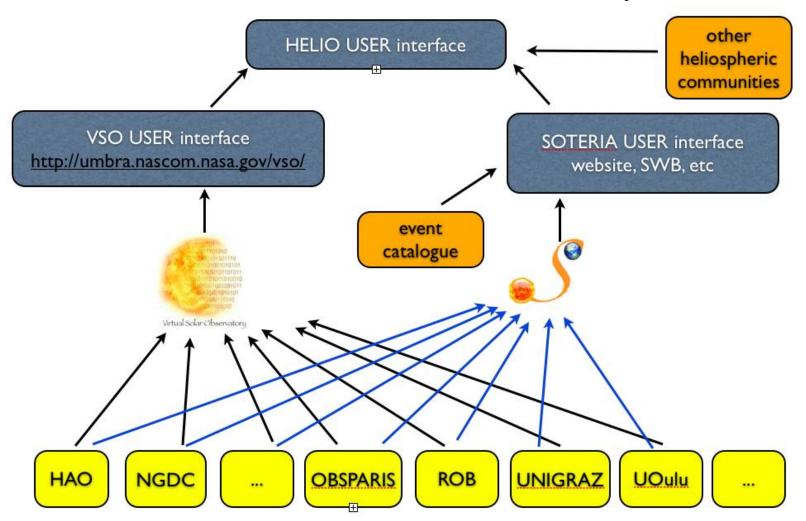


Expanding the global chemistryclimate-ionosphere model SOCOL-i to the lower ionosphere (Egorova et al., 2010)

#### **WP6: Successes**



### CASSIS FP7 coordination activity



## Major upcoming events



- School organised with COST and ICTP
- Capacity Building Workshop organised in Brussel
- Full exploitation of PROBA2 that now is operative
- Full exploitation of SDO that now is operative
- Focus on continuing and increasing the collaborations established so far
- Thematic workshops, visits and exchange of personnel among the beneficiaries
- Second annual meeting in October
- Web resources



### **Summer School**

#### **Summer School**

In collaboration with:

- COST Action on space weather
- ICTP Trieste
- To be held in Trieste
- Public outreach event, educational and involving scientists outside Soteria

# Capacity Building Workshop

Internal event organized by ROB Brussel

In Brussel

The scope is to train all Soteria members to the use of a common data management software and practice

# School on Space Weather

- Over 180 applicants
- Two weeks
- Speakers being invited now
- •Topics in space physics, space weather evets, effects on technology and people





# INTERNATIONAL ADVANCED SCHOOL ON SPACE WEATHER MODELLING AND APPLICATIONS

18 - 29 October 2010 Miramare, Trieste, Italy

The School is organised jointly by KCTP, the EC COST Action ES0803 "Developing Products and Sinvices for Space Weather in Europe", and the EC FP7 Project SOTERIA "SOLar-TERrostrial Investigations and Archives."

#### PURPOSE AND NATURE

Space Weather is the physical and phenomenological state of natural space environments under the effect of solar and non-solar driven perturbations.

In Europe the study of Space Weather has been promoted by the European Space Agency (ESA) and this triggered various cooperation initiatives like the torner EC COST Action 724 "Developing the Scientific Basis for Monitoring, Modelling and Predicting Space Weather", the new EC COST Action FSIMICK, and the FC FP7 Project SCITFRIA.

This school is a follow-up to the "International Advanced School on Space Weather" co-organized in 2006 by ICTP, EC COST Action 724, USNSWP, SCOSTEP/CAWSES, INAF, and INFN.

It is aimed at providing the scientific knowledge and the applied aspects of Space Weather, i.e., the monitoring and modelling resources based on advanced data handling, and if will address the following lopics: a. Space Weather Drivers and the Relevant Physical Environments; b. Space Weather Impacts on Technological Systems and Humans; c. Space Weather Monitoring and Data Handling: d. Space Weather Modelling Techniques. Morning sessions will be locused on theory and attempon sessions to practicals with the direct participation of the attendees.

#### **PARTICIPATION**

Scientists and students from all countries that are members of the United Nations, UNESCO or IAEA may attend the School. Although the main purpose of the ICTP is to help researchers from developing nations through a programme of training activities within a framework of international cooperation, students and postdoctoral scientists from developed countries are also welcome to attend. As the School will be conducted in English, participants must have a good working knowledge of that language.

As a rule, travel and subsistence expenses of the participants are borne by the home institution, However, limited funds are available for some participants (not more than 45 years of age) from, and working in, developing countries, to be selected by the organizers. Such financial support is available only for those who attend the entire activity. Every effort should be made by candidates to secure support for their fare (or at least half fare) from their home country. There is no registration fee to attend the School.

#### HOW TO APPLY FOR PARTICIPATION

The application form can be accessed at the activity website:

#### http://agenda.ictp.it/smr.php?2171

Once in the website, comprehensive instructions will guide you step-by-step, on how to fill out and submit the application form.

Telephone: +39-040-2240226 Telefax: +39-040-22407226

E-mail: smr2171@ictp.it ICTP Home Page: http://www.ictp.it

#### **DIRECTORS**

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- M. Messerotti (INAF, IT)
- G. Lapenta (KU Leuven, BE)
- S. Radicella (ICTP, IT)











#### MAIN CO-SPONSORS

- the Abdus Salam International Centre for Theoretical Physics (ICTP)
- EC COST Action E80803 "Developing Products and Services for Space Weather
- EC FP7 Project SOTERIA "SOLar-TERnistrial Investigations and Archives"
- National Institute for Astrophysics (INAF) (Italy)
- European Space Agency (ESA)

APPLICATION DEADLINE 31 May 2010